

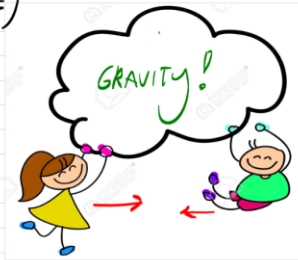
## Universal Gravitation and Orbits Examples

### EXAMPLES CONTINUED (FORCE)

$$\#1 \quad F = \frac{GMm}{r^2}$$

$$F = \frac{(6.67 \times 10^{-11})(80)(40)}{2^2}$$

$$F = 5.3 \times 10^{-8} \text{ N (VERY SMALL FORCE)}$$



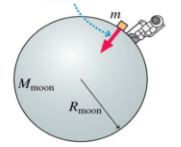
### GRAVITATIONAL FIELD STRENGTH

#1

$$g = \frac{GM}{r^2}$$

$$g = \frac{(6.67 \times 10^{-11})(7.35 \times 10^{22})}{(1.74 \times 10^6)^2}$$

$$g = 1.62 \text{ N/kg}$$



$$g = \text{--- m/s}^2$$

$$g = \text{--- N/kg}$$

$$\#2 \quad g = \frac{\text{FORCE}}{\text{MASS}} \quad \text{N/kg}$$

$$g = \frac{400 \text{ N}}{3 \text{ kg}}$$

$$g = 133.33 \text{ N/kg} \quad \text{or} \quad 133 \text{ m/s}^2$$



#3

$$g = \frac{GM}{r^2}$$

$$9.8 \text{ m/s}^2 = \frac{6.67 \times 10^{-11} (M)}{(6.38 \times 10^6)^2}$$

EXPERIMENT (pointing to 9.8 m/s<sup>2</sup>)  
EXPERIMENT (pointing to the equation)  
GREEKS (pointing to the denominator)

$$M = 5.98 \times 10^{24} \text{ kg}$$

$$\#4 \quad g = \frac{GM}{r^2}$$

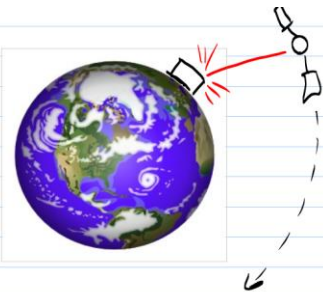
$$g = \frac{(6.67 \times 10^{-11})(5.98 \times 10^{24})}{(6.38 \times 10^6)^2}$$

$$g = 9.8 \text{ N/kg} \quad \text{or} \quad \text{m/s}^2$$

#5

Always simplify this equation before you begin to substitute numbers in

$$\frac{GMm}{r^2} = \frac{mV^2}{r}$$



$$\frac{GMm}{r^2} = \frac{mV^2}{r}$$

BEST TO SIMPLIFY FIRST  
(GETS RID OF SOME BIG NUMBERS)

①

$$\frac{GM}{r} = V^2$$

(CANCEL "m" AND MULTIPLY BOTH SIDES BY "r")

②

$$V = \sqrt{\frac{GM}{r}}$$

$$\leftarrow r = (4.4 \times 10^5 + 6.38 \times 10^6 \text{ m})$$

↑  
ABOVE  
EARTH

↑  
RADIUS  
OF EARTH

③

$$V = 7647 \text{ m/s}$$

#6

$$\frac{GMm}{r^2} = \frac{4m\pi^2 r}{T^2}$$



$$\textcircled{1} \quad \frac{GMm}{r^2} \left(\frac{1}{r}\right) = \frac{4m\pi^2 r}{T^2} \left(\frac{1}{r}\right)$$

(GET RID OF  $m$  +  $r$   
FIRST... THEN SUB  
IN NUMBERS)

$$\textcircled{2} \quad \frac{GM\cancel{m}}{r^3} = \frac{4\cancel{m}\pi^2 r}{T^2}$$

$$\textcircled{3} \quad \frac{GM}{r^3} = \frac{4\pi^2}{T^2}$$

$$\textcircled{4} \quad \frac{1}{r^3} = \frac{4\pi^2}{GMT^2}$$

$$\textcircled{5} \quad r^3 = \frac{GMT^2}{4\pi^2}$$

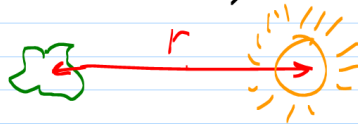
$T$  - PERIOD MUST BE IN SECONDS

$$T = 187 \times 24 \times 60 \times 60$$

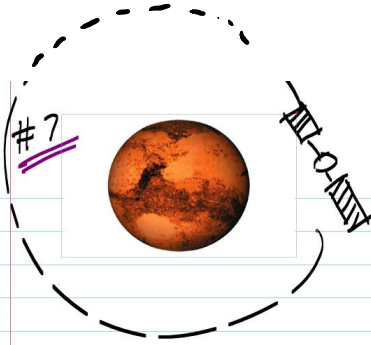
$$T = 1.616 \times 10^7$$

$$r = 8.7 \times 10^{10} \text{ m}$$

(DISTANCE BETWEEN CENTRE OF  
KOKO AND IT'S STAR.)



#7



$$\frac{GMm}{r^2} = \frac{mV^2}{r}$$

$$\frac{GM}{r^2} = \frac{V^2}{r}$$

$$\frac{GM}{r} = V^2$$

Simplify, THEN SUB

$$r = \frac{GM}{V^2}$$

$$r = \frac{(6.67 \times 10^{-11})(6.37 \times 10^{23} \text{ kg})}{(2000)^2}$$

$$a) \quad r = 1.062 \times 10^7 \text{ m}$$

b) RADIUS OF MARS (INTERNET)  
 ( $3.39 \times 10^6 \text{ m}$ )

$$\rightarrow 1.062 \times 10^7 \text{ m} - 3.39 \times 10^6 \text{ m} = 7.23 \times 10^6 \text{ m (ABOVE MARS)}$$